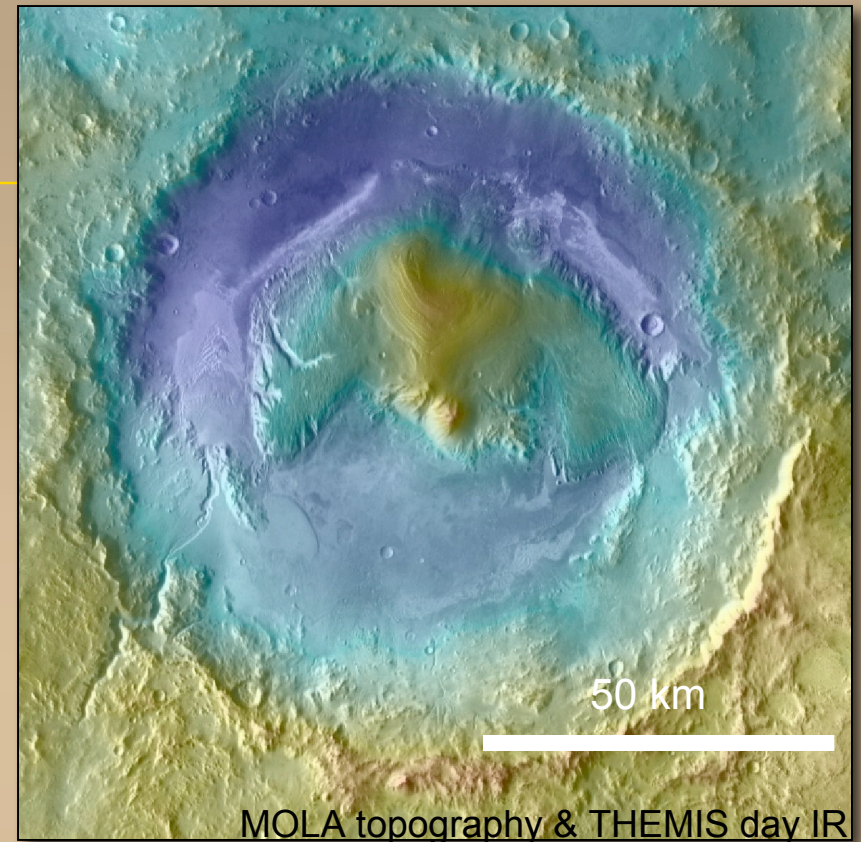
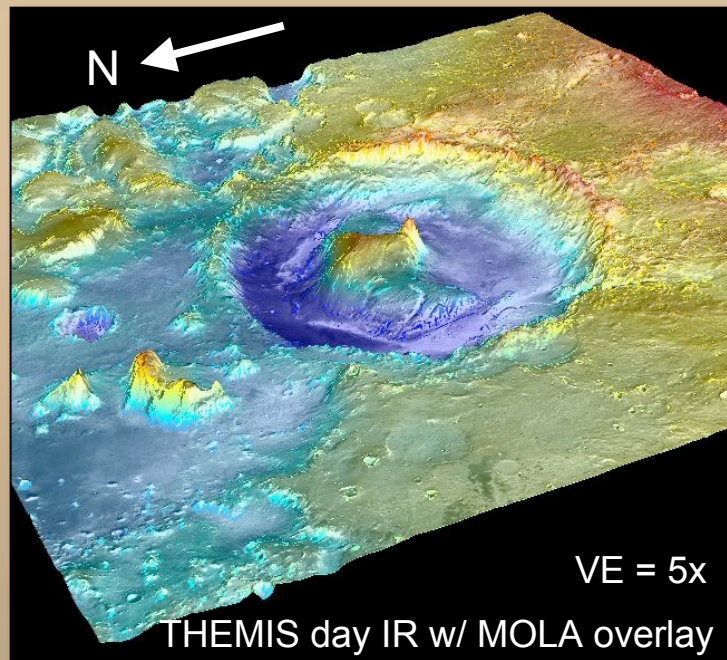


Gale Crater: Context and layer diversity from HiRISE images



Brad Thomson
Nathan Bridges

JPL/Caltech

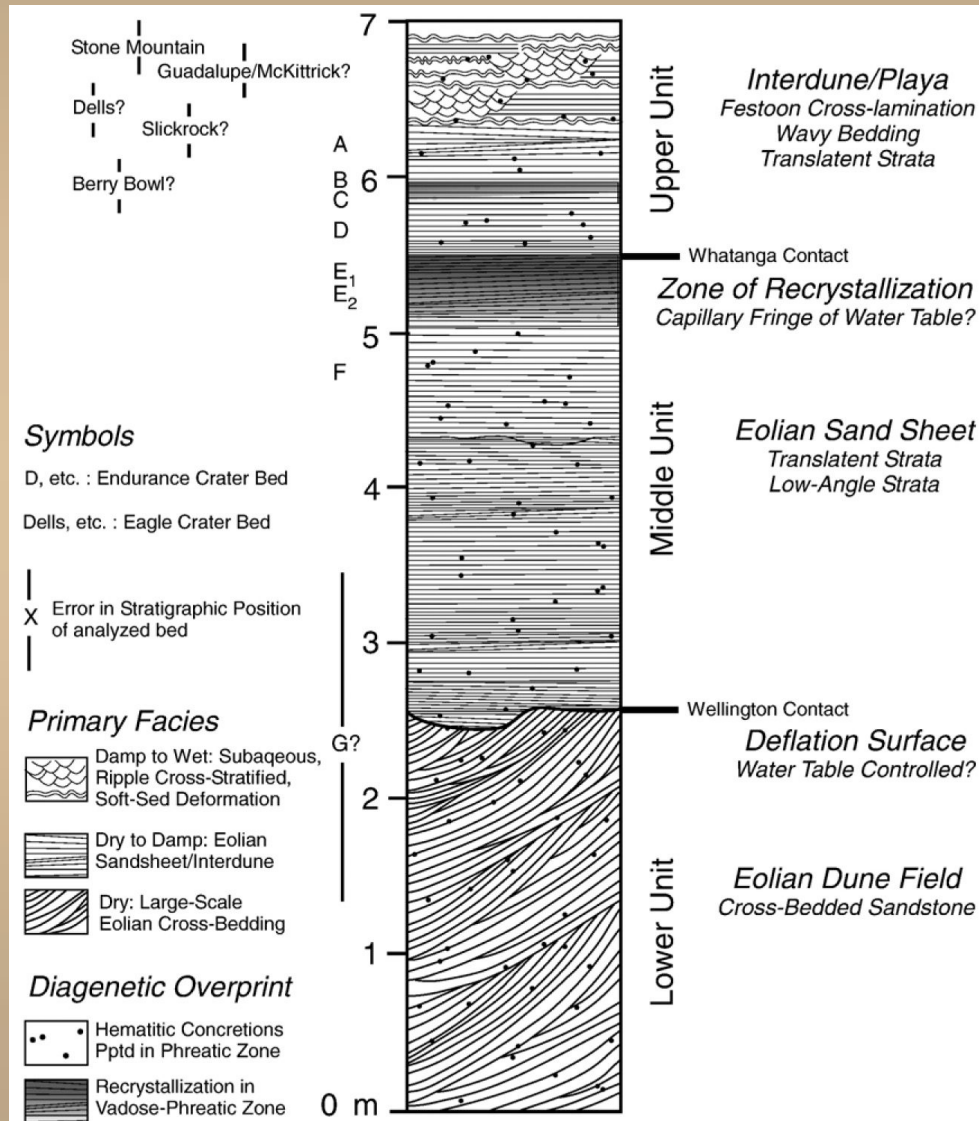
Key Moon-Mars differences



1. Outflow channels
 - VL1, Pathfinder
2. Valley networks
 - Spirit
3. Finely layered sedimentary sequences
 - ~Opportunity, MSL?

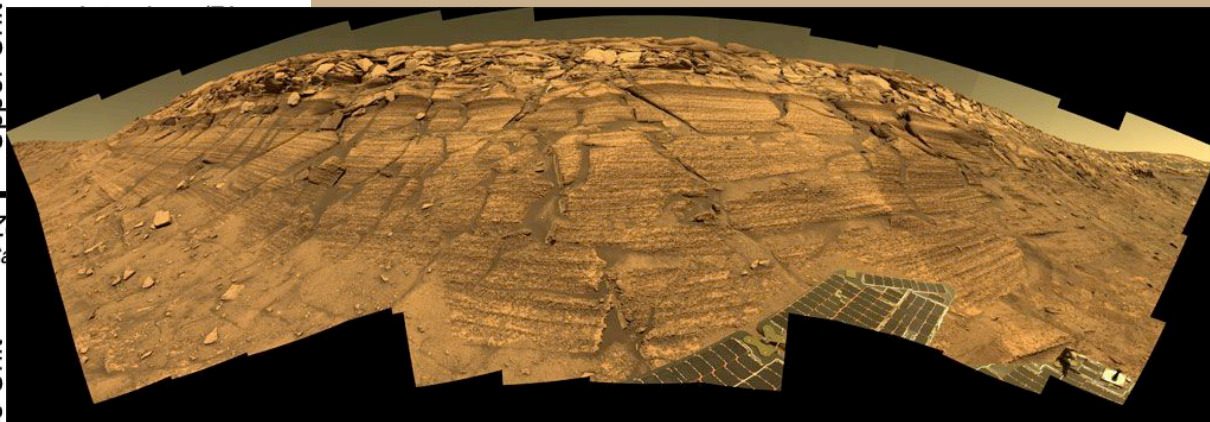
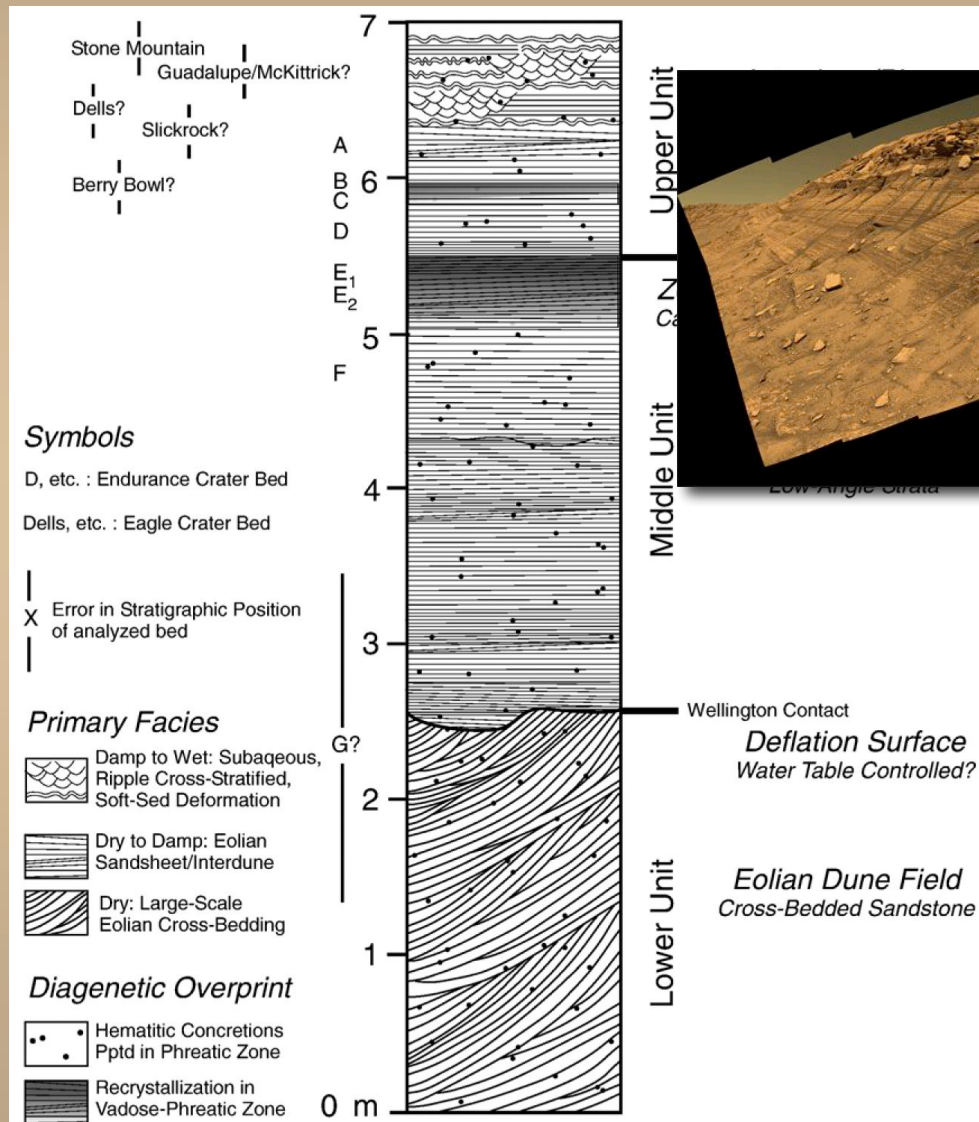


Stratigraphic Framework



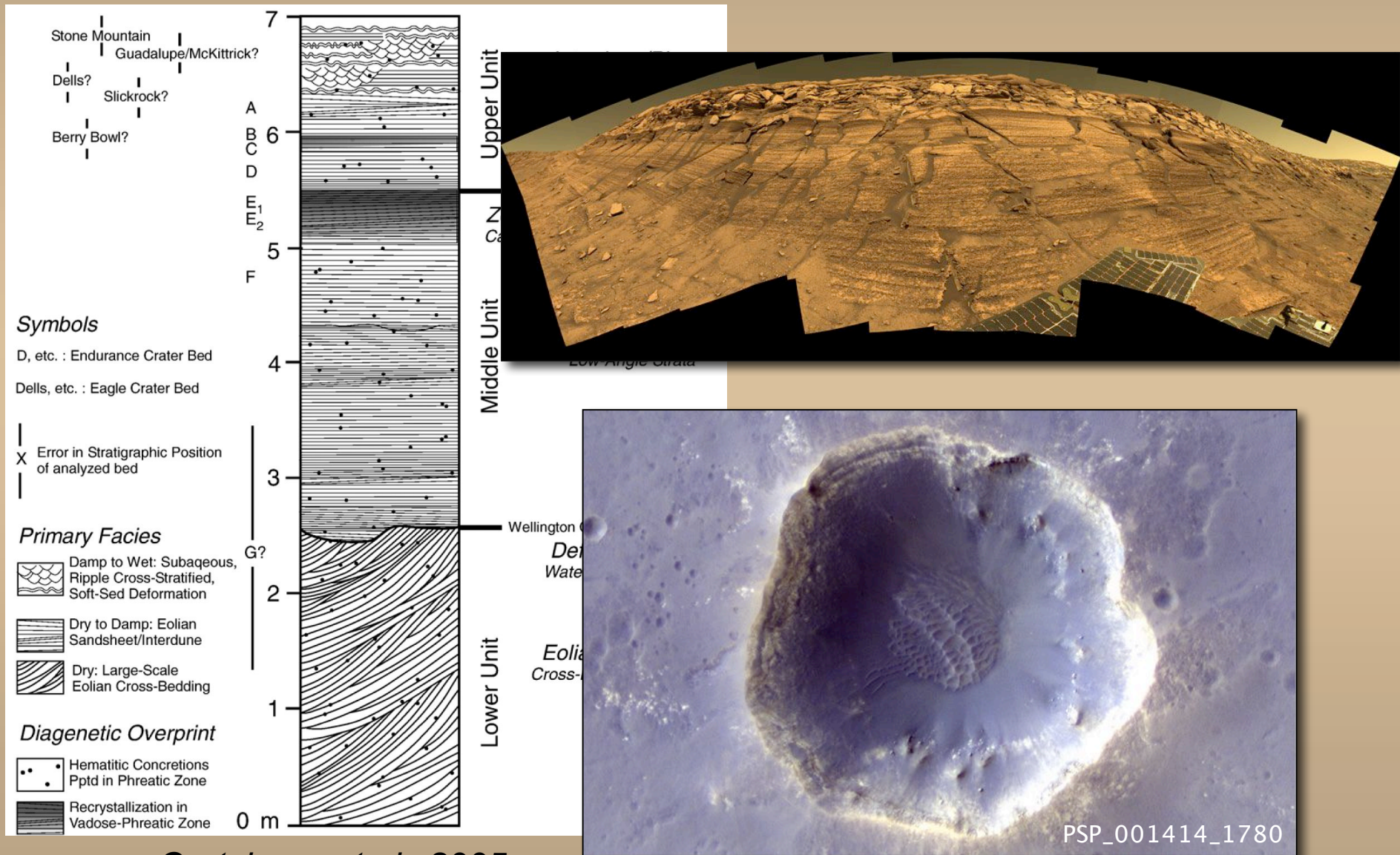
Grotzinger et al., 2005

Stratigraphic Framework

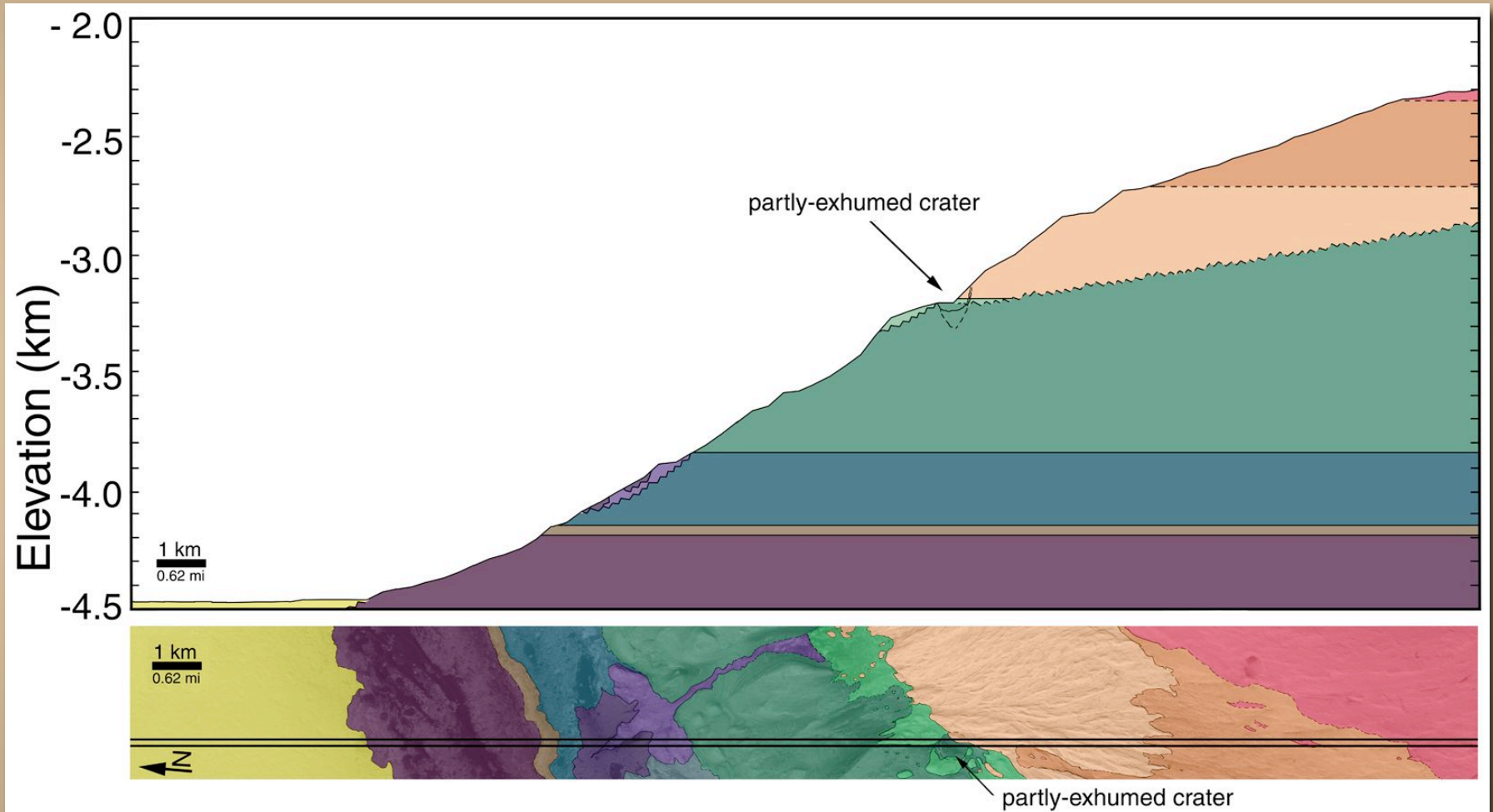


Grotzinger et al., 2005

Stratigraphic Framework



Geologic cross-section

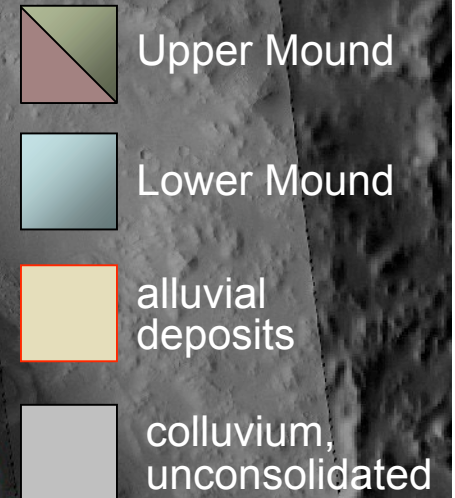


V.E. = 4.7

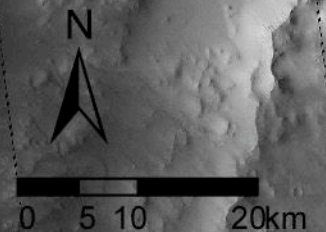
*Edgett & Malin, 2000 LPSC
Malin & Edgett, 2000 Science*

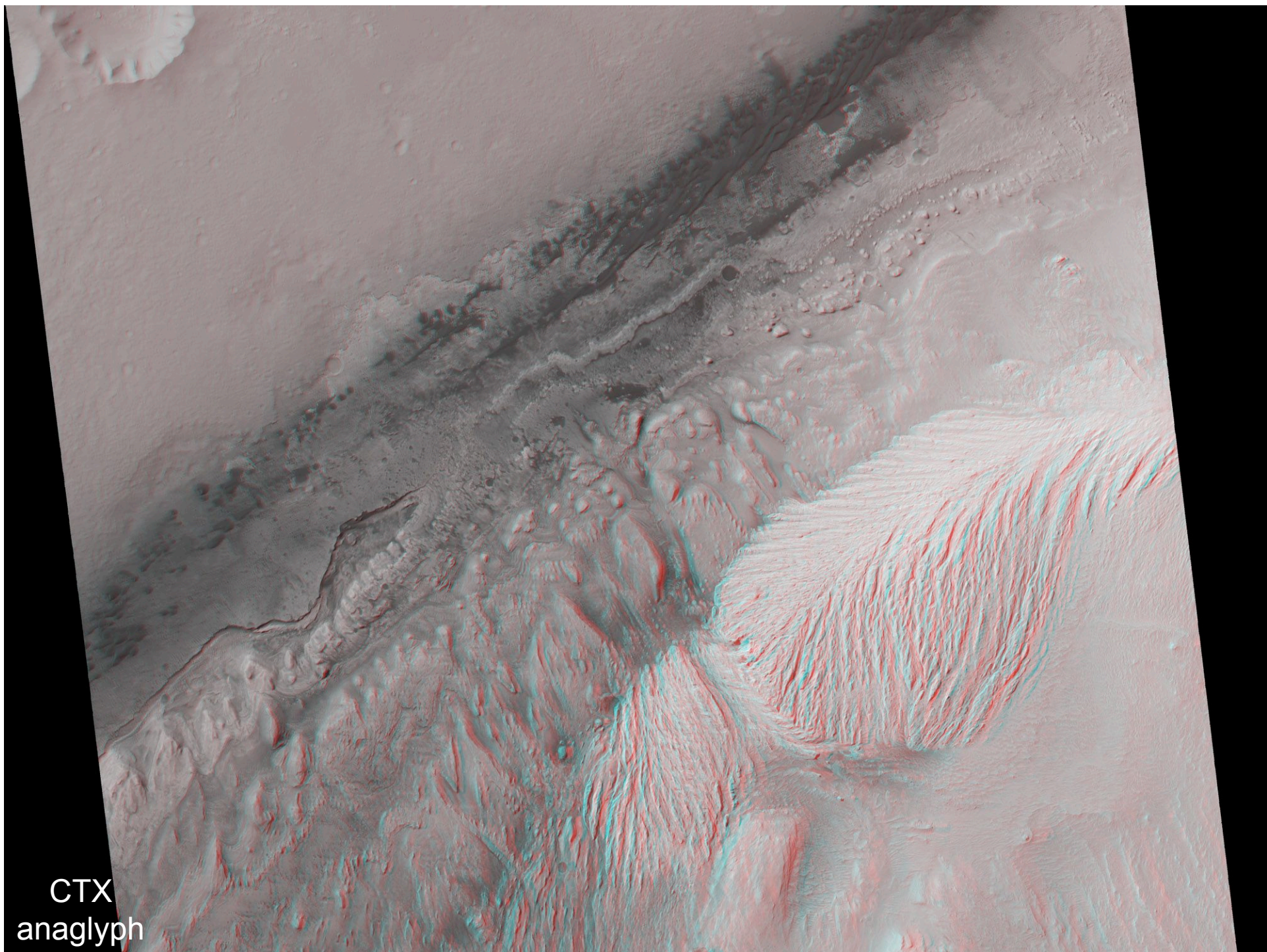
Gale HiRISE/CTX geomap

Simplified key



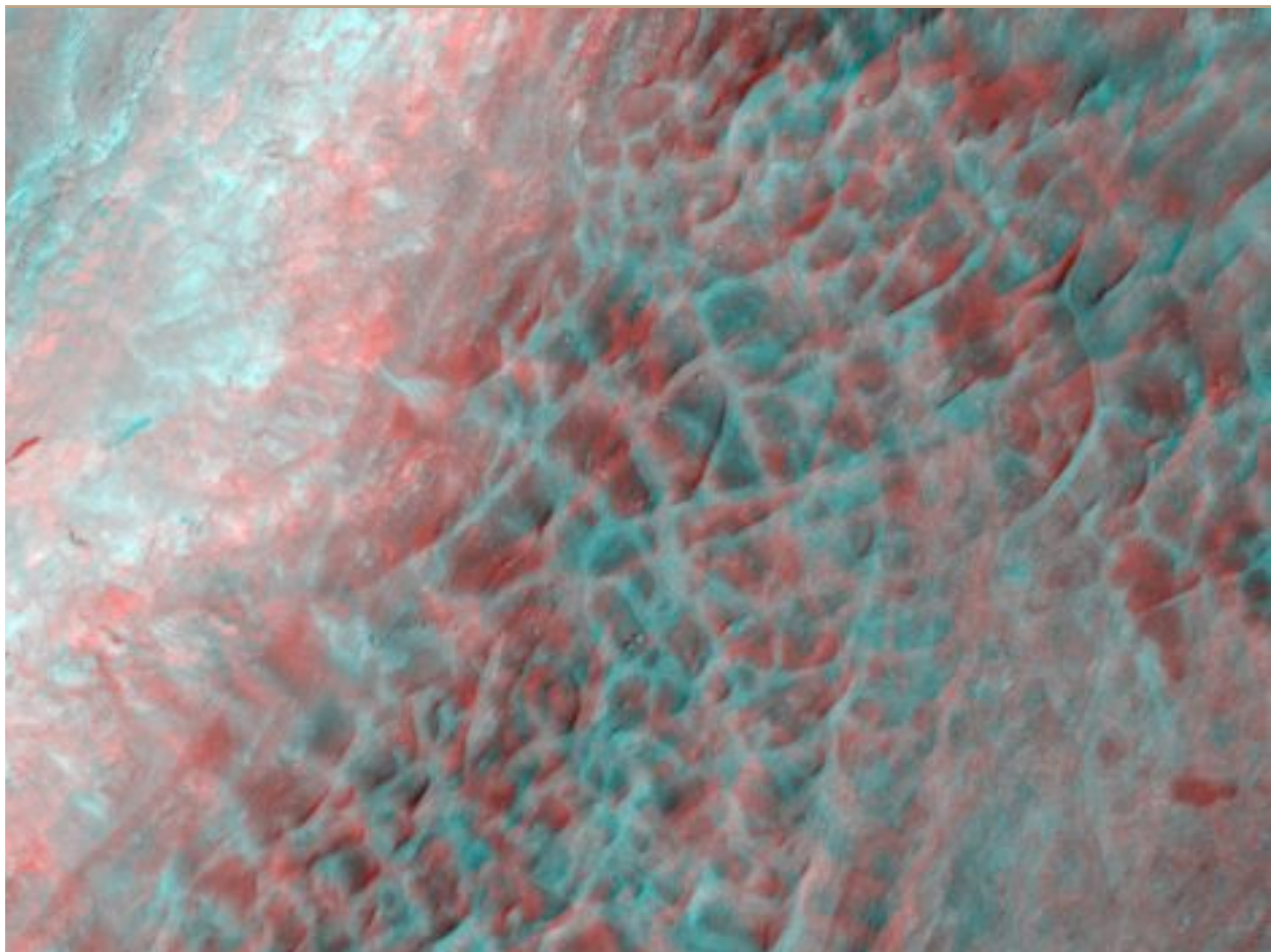
CTX mosaic

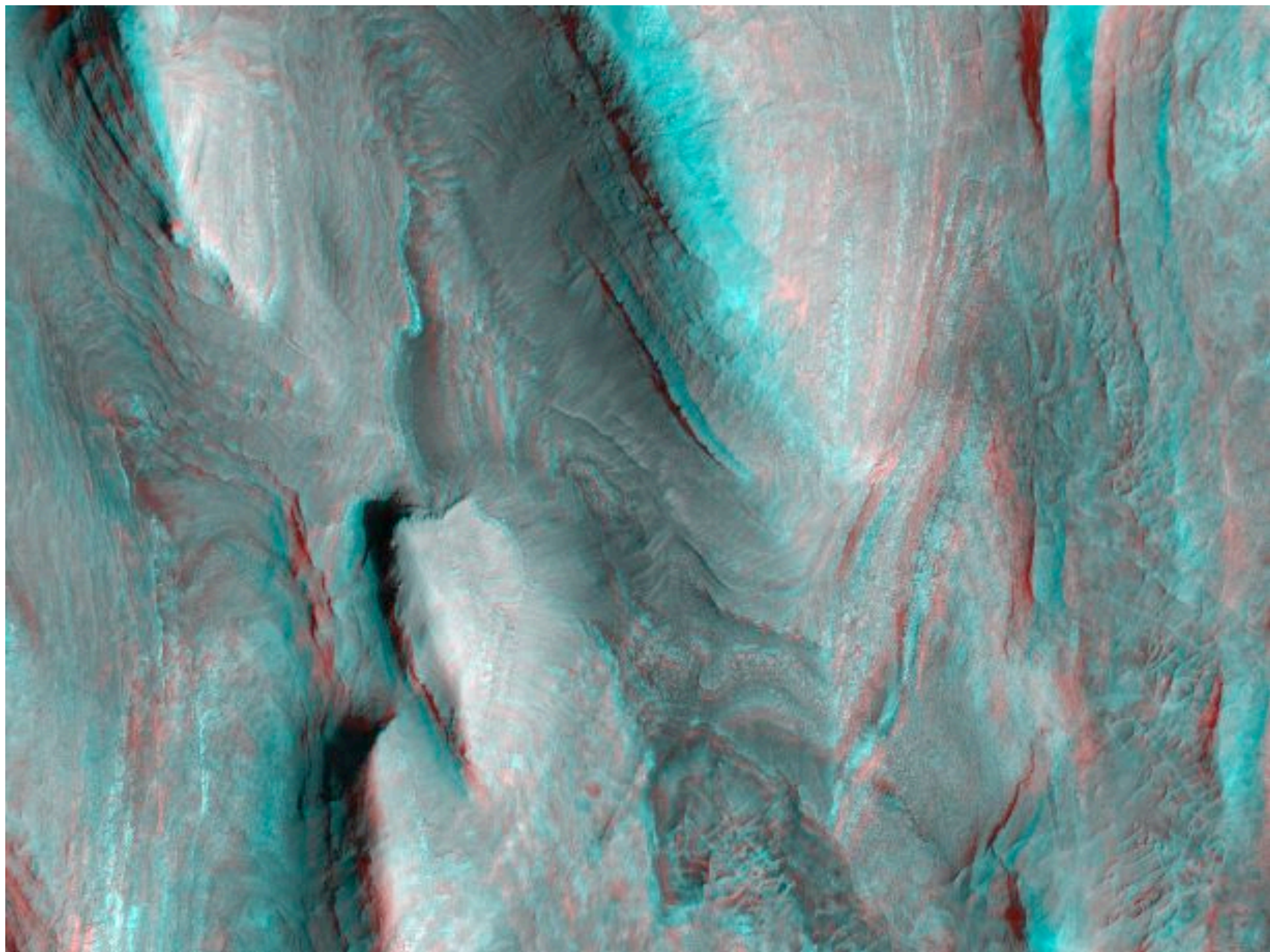




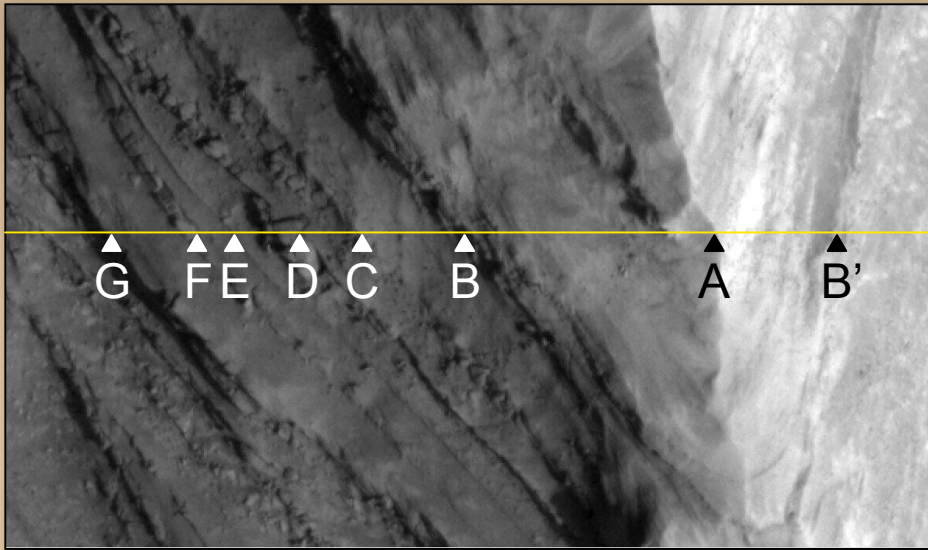
CTX
anaglyph

Lower Mound



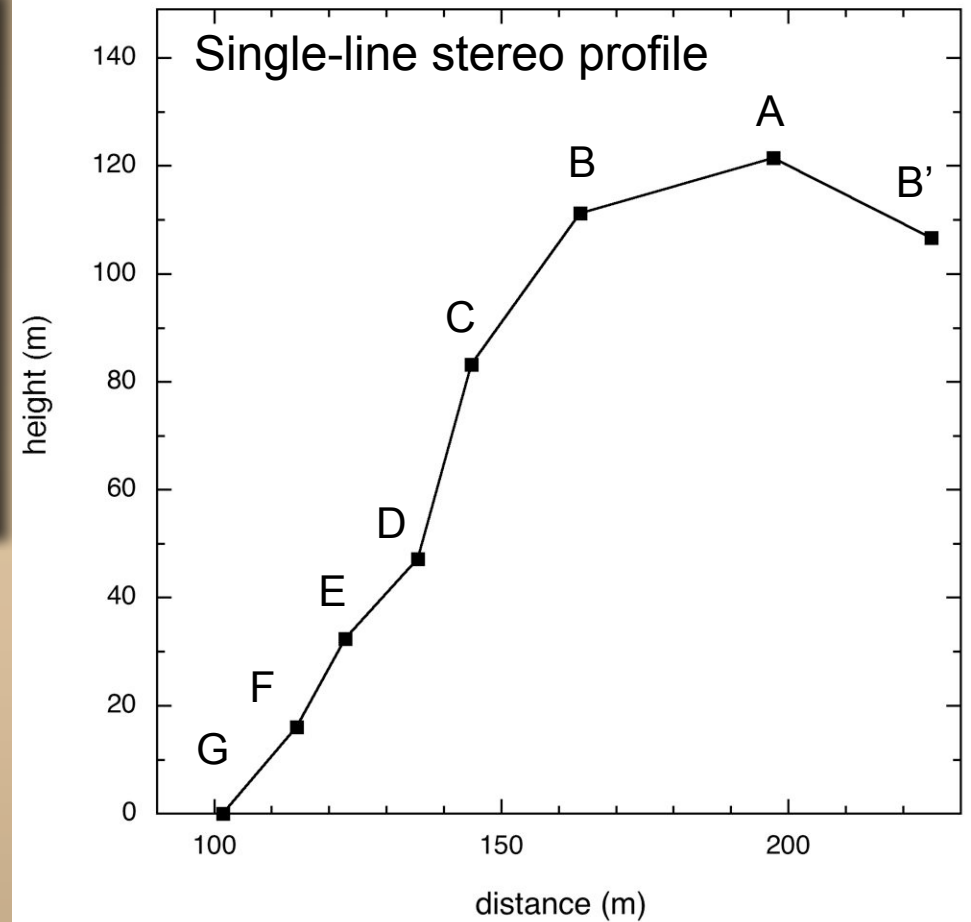


Lower mound layers

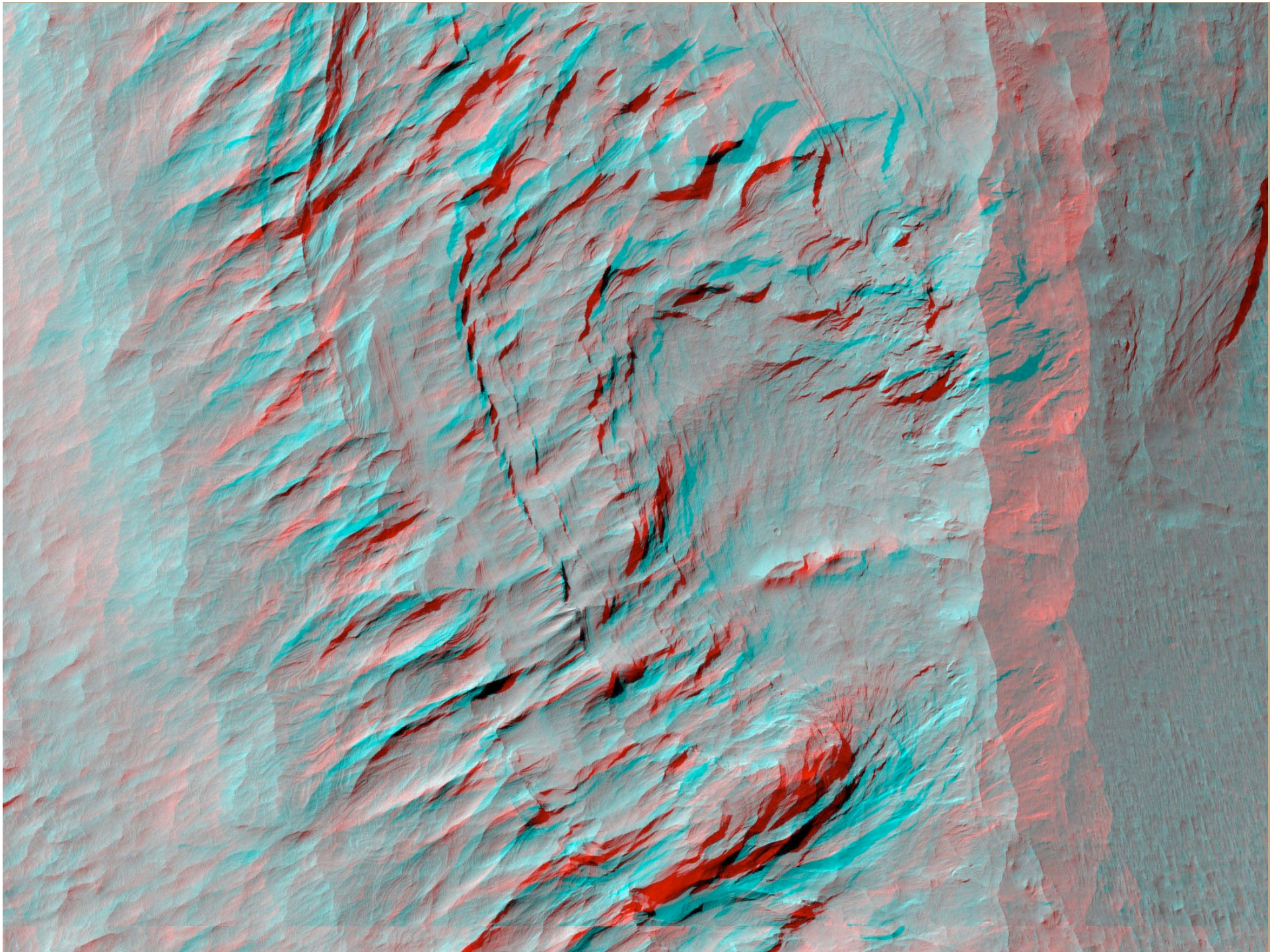


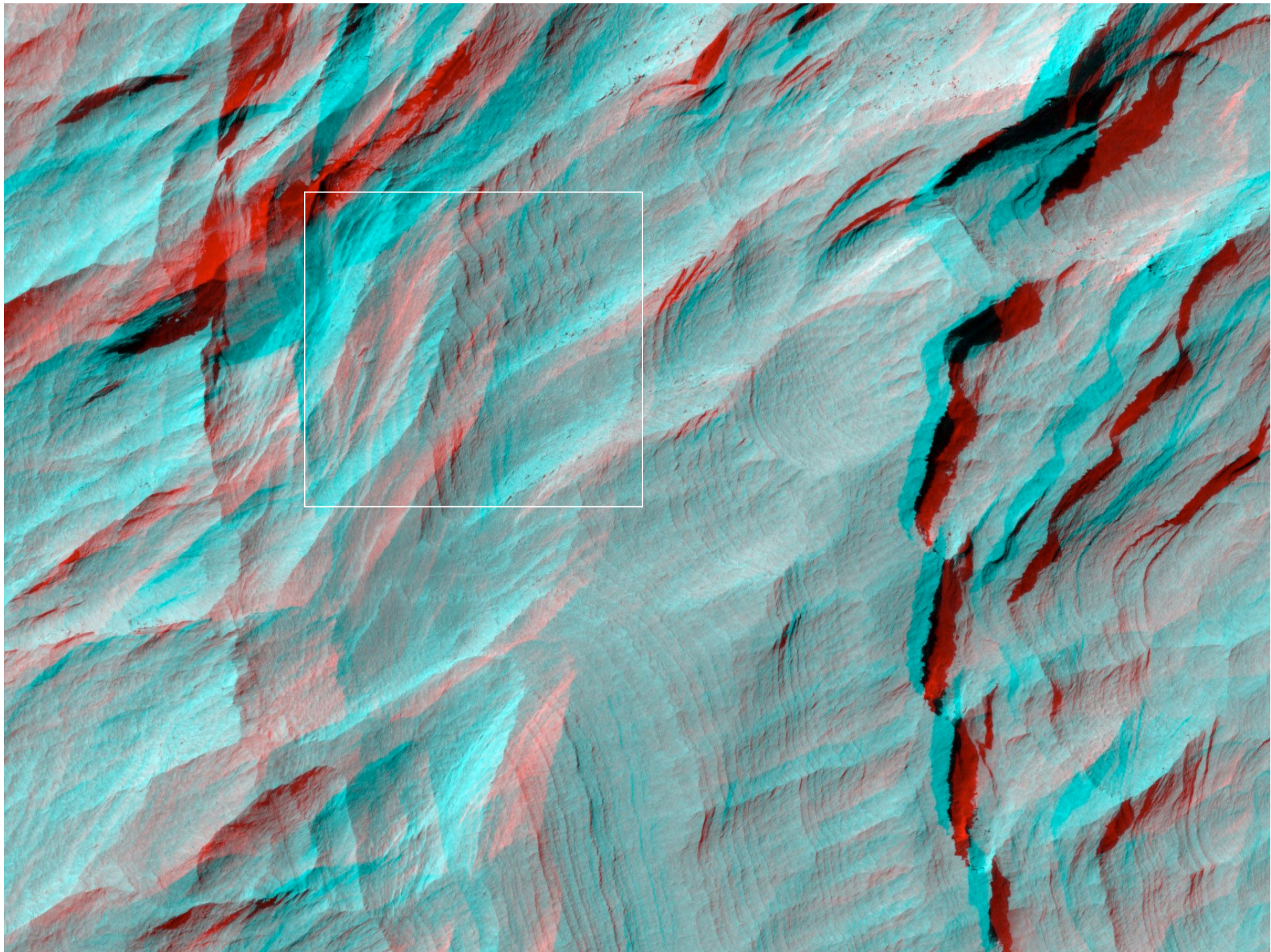
PSP_001488_1750
(& PSP_001752_1750)

- Layers 10-30 m thick
- Apparent slope $\sim 2^\circ$



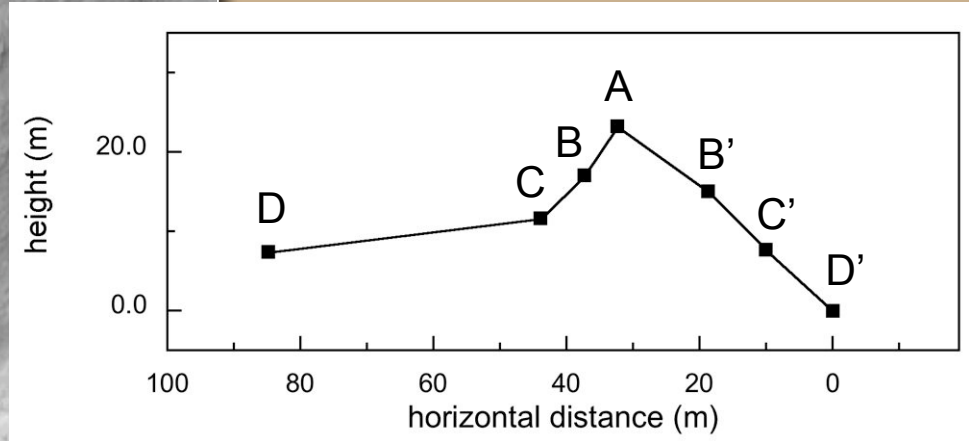
Upper Mound





Upper mound layers

Single-line stereo profile

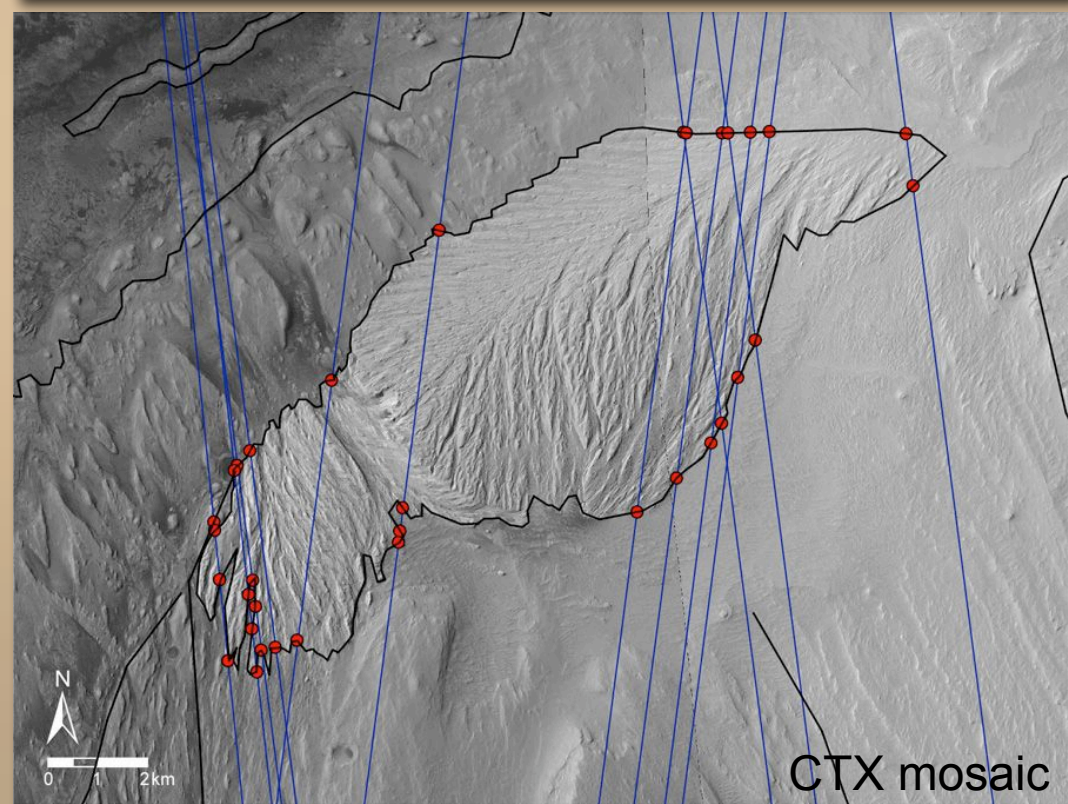
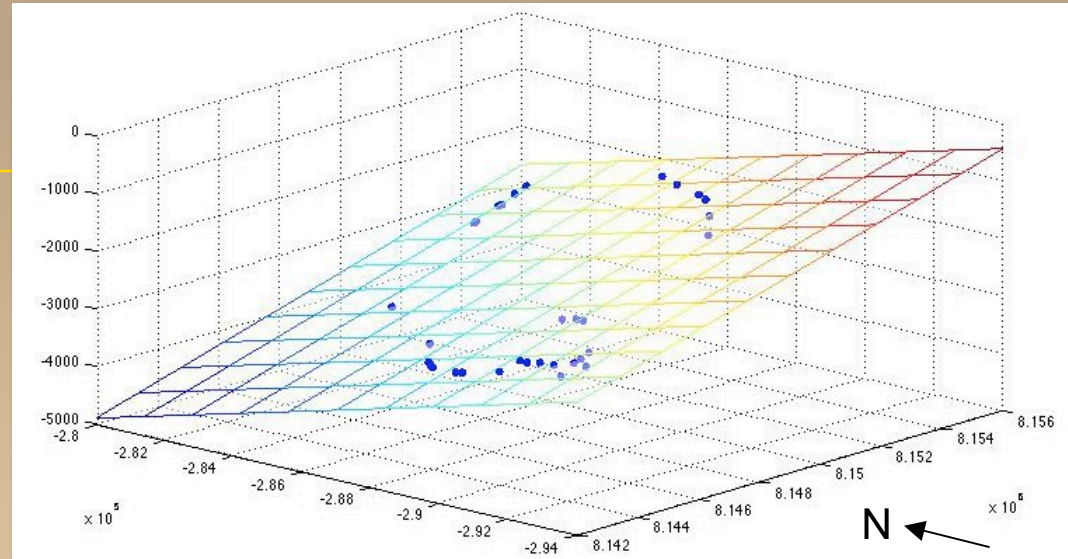


- Layers ~3-7 m thick
- Apparent slopes s_a :
 $5^\circ > s_a < 10^\circ$
- Dip azimuth is to east (limited to 2D)

PSP_009927_1750 & (PSP_008002_1750)

Contact geometry

- Mapped upper unit contact (unconformity)
- Extracted MOLA interpolated elevation points along contact
- Best-fit plane is non-horizontal
- Geometry inconsistent with simple lacustrine depositional process
 - Max elevation difference ~1.6 km
 - Suggestive of eolian control of surfaces of omission

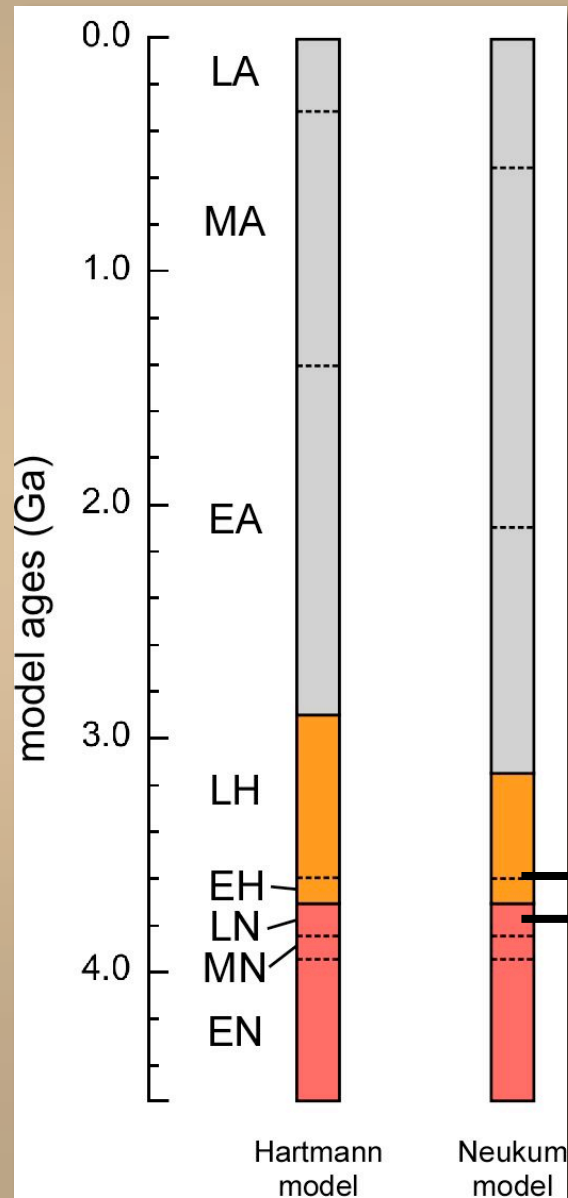


Diversity summary

Lower Mound units have distinct characteristics from Upper Mound, suggesting environmental differences

- Lower Mound:
 - Units with distinct mineralogic signatures that result from aqueous alteration
 - Cemented and inverted fractures suggests fluid flow (potentially indicates some in situ alteration)
- Upper Mound:
 - Erosional morphology, thermal inertia suggests fine-grained component, possibly eolian
 - Attitude of erosional unconformity suggests eolian control of surfaces of omission
 - Channels sourced from Upper Mound suggest additional volatile components

Age context



- Continuing erosion
- Onlapping valley network deposition on crater floor
- Interior channel deposits
- Deposition of upper mound layers
- Depositional hiatus / erosional episode
- Deposition of lower mound layers
- Impact of Gale Crater-forming bolide

Hartman & Neukum, 2001

Overall summary

Context: Low-energy sedimentary depositional environment; general age constraints can also be inferred.

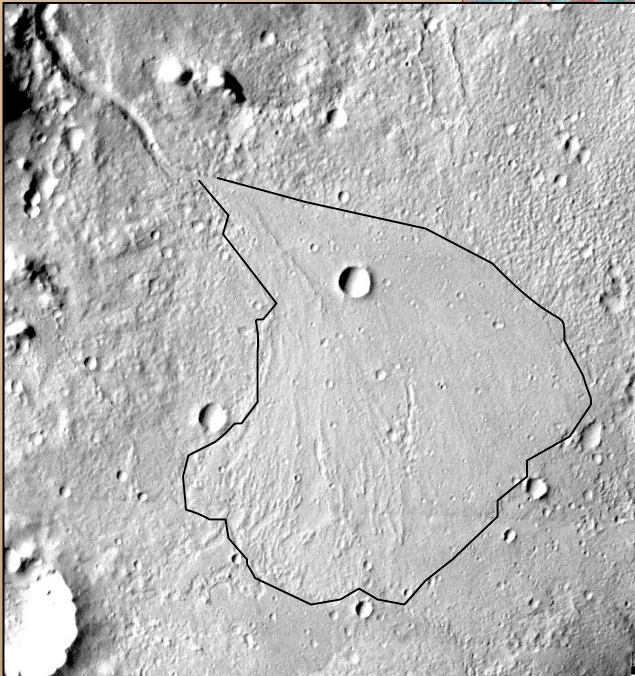
- Closed depositional basin.
- Gale representative of numerous exposures of layered sedimentary units on Mars, non-uniform in space and time
- Long sedimentary record captures environmental changes during deposition.
- Formation mechanism: What can we rule out?
 - **Unlikely**: impact ejecta, effusive volcanism, pedogenesis (acting alone), volcaniclastics (lack of regular repetition)
 - **Possible**: lacustrine deposition, eolian deposition

Context summary

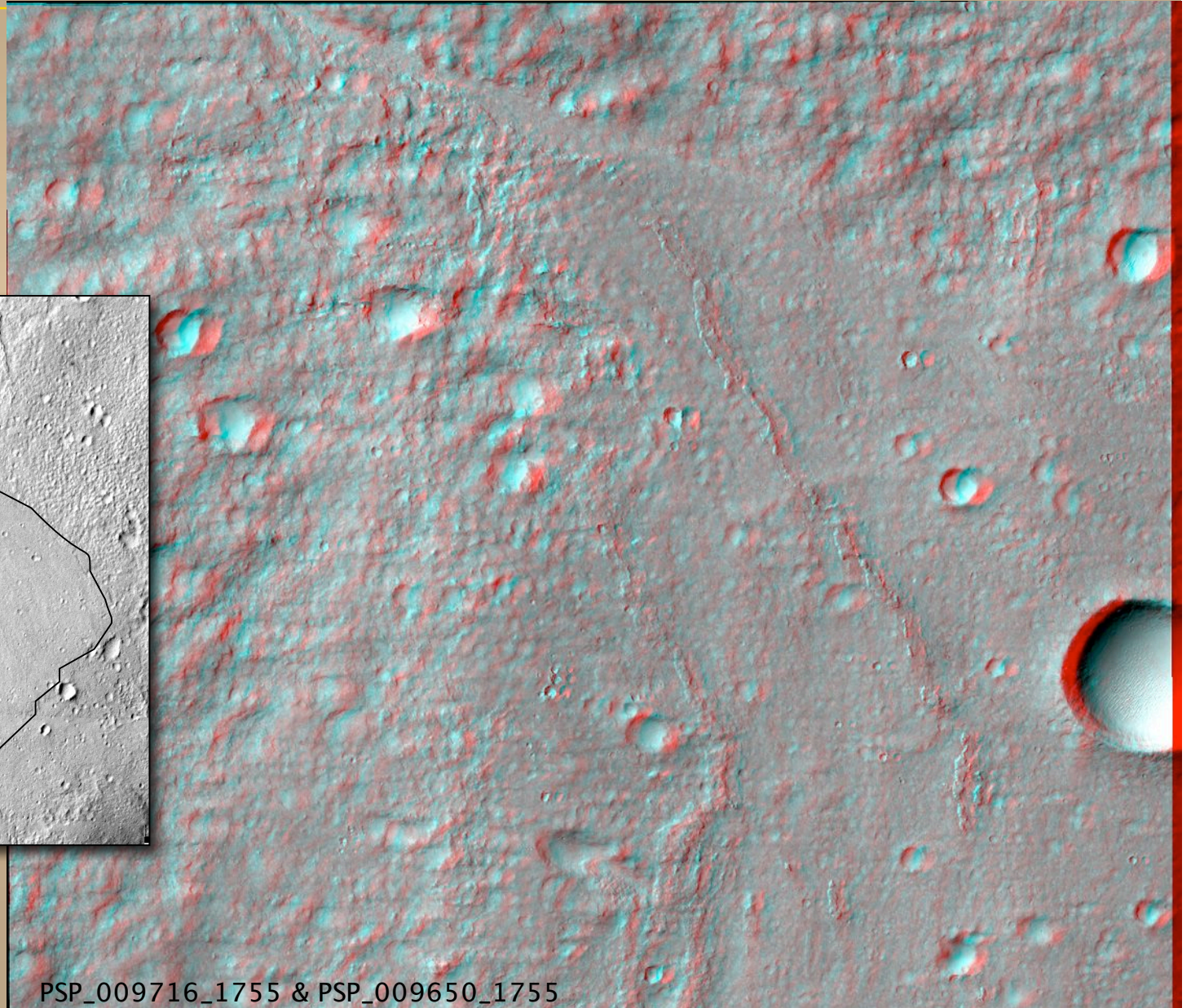
Context: Low-energy sedimentary depositional environment; general age constraints can also be inferred.

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 - Long sedimentary record captures environmental changes during deposition.
 - Formation mechanism: What can we rule out?
 - **Unlikely**: impact ejecta, effusive volcanism, pedogenesis (acting alone), volcaniclastics (lack of regular repetition)
 - **Possible**: lacustrine deposition, eolian deposition
- Low-energy deposition of particles via settling from suspension in a fluid (wind or water). Even if eolian, aqueous processes played a significant role.

Extra: Ellipse science



THEMIS vis



PSP_009716_1755 & PSP_009650_1755